PARTICLE-BASED SIMULATIONS OF MICROSCOPI
THERMAL PROPERTIES OF CONFINED SYSTEMS

Flavio Sabatti, Ky Merrill, Stephen Goodnick, and Marco Saraniti
Arizona State University

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Final Report

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1.0 Summary

In this document, we offer the final report of the Defense Advanced Research Projects Agency (DARPA)/United States Air Force (USAF) award # FA8650-10-1-7045. An overview of the project is discussed, and the progress of the research activity is assessed, as related to the program milestones.
2.0 Introduction

Within this project, a novel modeling approach is being defined for the development of a transformational technology based on the microscopic thermal management of solid-state devices. By the full inclusion of the phonon dynamics within the framework of charge transport, we are implementing microscopic models in our existing particle-based Cellular Monte Carlo (CMC) computer-aided design (CAD) tools [1] for the design of electron devices where the generation and the transport of thermal energy are functionally coupled with their electrical specifications. The crucial challenges of the present phase of the project are summarized in the following three tasks:

1. Implementation of a rejection algorithm for the electron-phonon scattering table.
2. Realization of a solver for the heat transport equation.
3. Integration of phonons and electrons in a unified particle-based framework.

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Accordingly to the statement of work, such tasks have been organized in the following milestones:

This is the final report that summarizes the methodology developed to model phonon-phonon and phonon-electron interaction within a perturbative full-band approach.
3.0 Methods, Assumptions, and Procedures

3.1 Rejection Algorithm for Scattering

The capabilities of the original CMC algorithm [1] have been extended using a rejection algorithm. This approach retains the simulation speed advantages of the CMC scheme while allowing the adaptation of the scattering rates to the real time local conditions. Figure 1 shows the rejection algorithm flow chart. After the CMC triggers an event with probability $P_{CMC}$, a new scattering probability $P_{loc}$ is computed from the local conditions. The rejection probability $P_{rej}=P_{loc}/P_{CMC}$ is then used in a stochastic procedure to decide whether the scattering event occurs or is rejected.

The scattering probability after the rejection is

$$P_{q\rightarrow q'} = P_{CMC} \times \left( \frac{P_{loc}}{P_{CMC}} \right) = P_{loc}$$  \hspace{1cm} (1)

where $q$ is the initial state, and $q'$ the final state.

![Figure 1: Flow Chart of the Rejection Algorithm](image-url)
3.2 Phonon-Phonon Scattering Rate

Perturbation theory is used to implement both phonon-phonon and phonon-defect interactions. The anharmonic decay and recombination scattering rate based on the Klemens model [2] is given by the following expression:

\[ P_{q-q'} = \frac{2\hbar}{M^3} \frac{c^2(q, q', q'')}{\omega \omega' \omega''} \delta(\omega, \omega', \omega'') \delta(q, q', q'') F_{\text{loc}} \]  

where \( F_{\text{loc}} = F(n_{\text{loc}}', n_{\text{loc}}'') \) is a local population dependent factor, \( n_{\text{loc}}' \) and \( n_{\text{loc}}'' \) are the populations at \( q' \) and \( q'' \), respectively, \( M \) is the average mass, \( c^2(q, q', q'') \) is the anharmonic coefficient obtained from the perturbation Hamiltonian, and \( \omega, \omega' \) and \( \omega'' \) are the angular frequencies corresponding to the states \( q, q' \), and \( q'' \), respectively.

The phonon-defect scattering probability based on the work of Srivastava [3] is given by:

\[ P_{q-q'} = \frac{V_w \Gamma}{6 \omega \omega' (e_q \cdot e_{q'})^2} \delta(\omega - \omega')(n_{\text{loc}}' + 1) \]  

where \( e \) is the polarization vector, \( V_w \) is the Wigner-Seitz cell volume, and \( \Gamma \) is an expression that depends on the phonon-defect type to be discussed later.

Computing the local scattering probability requires estimating the local phonon populations \( n_{\text{loc}}' \) and \( n_{\text{loc}}'' \) from the ensemble. This task requires defining a sampling volume of the particle position both in real space \( (V_{\text{loc}}) \) and in momentum space \( (\Omega_{\text{loc}}) \), then counting the number of simulated phonons \( \eta(\Omega_{\text{loc}}, V_{\text{loc}}) \) in the volume, and, finally, obtaining the phonon population using the following approximation:

\[ n_{\text{loc}} = \frac{\eta(\Omega_{\text{loc}}, V_{\text{loc}})}{N_{\text{state}}} = \frac{\eta(\Omega_{\text{loc}}, V_{\text{loc}}) \times 8\pi^3}{\Omega_{\text{loc}} V_{\text{loc}}} \]  

where \( N_{\text{state}} \) is the number of states in the volume.

3.3 Phonons Initialization and Boundary Conditions

The number of equilibrium phonons within a sampling volume \( V_{\text{loc}} \) in a crystal of volume \( V_c \) is:

\[ N_{\text{pho}}(T, V_{\text{loc}}) = \frac{V_{\text{loc}}}{V_c} \sum_p \sum_q \langle n_{q,p}, T \rangle \]  

where \( p \) is the phonon mode index, \( \langle n_{q,p}, T \rangle \) is the Bose-Einstein distribution, and \( T \) is the temperature. Typically, \( N_{\text{pho}}(T, V_{\text{loc}}) \) is an extremely large number, and even within the relatively small volume of a modern transistor there are too many particles to simulate individual phonons practically. For this reason, a weighting factor \( W \) is used, and each simulated particle represents \( W \) real particles. The factor \( W \) is computed before starting the simulation using the formula:

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\[ W = \frac{N_{\text{dev}}}{N_{\text{sim}}}, \]  

where \( N_{\text{dev}} \) is the total number of real phonons in the device and \( N_{\text{sim}} \) the number of super particles to be simulated.

The particle-based phonon Monte Carlo code implements three boundary conditions: (1) planar interfaces, (2) rough interfaces, and (3) contacts. The first boundary condition models ideal surfaces, in which the incident particle is specularly reflected. The second boundary condition models interfaces between the extreme cases of perfectly specular reflection, considered above and perfectly diffuse reflection. The model is based on the work of Fuchs-Sondheimer [4] [5] and adapted to the full band approach. Contacts are modeled as ideal reservoirs with infinite thermal capacity and constant temperature. The contact population is replenished injecting phonons in the reservoir using a velocity-weighted probability distribution, similar to Pardo [6].
4.0 Results and Discussion

4.1 Validation of the Phonon Scattering Rate

Tests have been designed to verify the algorithm’s capability of adapting to the real-time local particle distribution. The first test verifies the capability of the scattering algorithm to reach the equilibrium steady state distribution, while a second test verifies the capability of reproducing experimentally observed results over a wide range of temperatures.

Let’s consider an adiabatic system with total energy density $E_c$. For any initial energy distribution, the phonons will eventually reach a Bose-Einstein distribution with temperature $T_c$ ($\langle n_{q,p}, T_c \rangle$) that satisfies the following energy vs. temperature relationship:

$$E_c(T_c) = \frac{1}{V_c} \sum_p \sum_q \left( \langle n_{q,p}, T_c \rangle + \frac{1}{2} \right) \hbar \omega_q$$

Figure 2 shows the energy distribution at the initial and the final state of the simulation. The system is initialized with energy density corresponding to a temperature of 200 Kelvin (K), $E_c(200K)$, but the initial distribution is set as that corresponding to a temperature of 300K, $d_{\text{ini}}(q, p) \propto \langle n_{q,p}, 300 \rangle$ (red line), The system is allowed to evolve until it reaches steady state (black line), at which time the energy density corresponding to the theoretical distribution $\langle n_{q,p}, 200K \rangle$ (blue line) is obtained.

![Figure 2: Simulated Evolution of a System Initialized with a Non-equilibrium Distribution of Phonons](image)

*The system evolves to a final distribution coincident to the analytical reference distribution.*
4.2 Phonon-Isotope Scattering

An isotope is an atomic variant of an element with the nominal number of protons, but a different number of neutrons and hence a different atomic mass. For example, $C^{12}$, $C^{13}$, and $C^{14}$ are isotopes of Carbon having 12, 13, and 14 neutrons, respectively. The isotopes of an element show identical chemical behavior (except in the reaction speed) as well as both crystal and electronic structure, however, the difference in mass disrupts the periodicity of the crystal resulting in a large phonon scattering rate. The overall effect of this is seen in a large reduction of the thermal conductivity tensor ($k$).

The thermal conductivity of the natural crystal ($k_{nat}$) and the isotopically pure crystal ($k_{iso}$) can differ by up to one order of magnitude at low temperatures [7]. At room temperature the effect is more ambiguous, and even for a material studied as extensively as silicon (Si) the reported $k_{iso}$ range is between 110% [8] [9] and 160% [10] [11] [12] of the $k_{nat}$ value. For these reasons isotope scattering cannot be overlooked in the simulation of real semiconductor across different ranges of temperature. Isotope scattering is implemented as a phonon-defect scatter; using Srivastava’s model expressed in equation (3) where the parameter $\Gamma$ is given by:

$$\Gamma = \sum_i f_i \left(\frac{\Delta M_i}{\bar{M}}\right)$$  \hspace{1cm} (8)

and where $f_i$ is the fraction of atom with mass $M_i$, and $\bar{M}$ the average mass.

Figure 3 shows the effect on the thermal conductivity of Si of the scattering model implemented in our particle based code. The lines represent the measured [8] thermal conductivity of natural silicon $Si^{nat}$ (92.2% $Si^{28}$, 4.6% $Si^{29}$ 3.1% $Si^{32}$) and an isotopically enriched sample of $Si^{28}$. The simulations without the isotope scattering follow the isotopically pure thermal conductivity, while including the isotope scattering model allows for accurate reproduction of the $Si^{nat}$ thermal conductivity.
Figure 3: Simulated and Experimentally Determined Effects of Phonon-Isotope Scattering on Low Temperature Thermal Conductivity: the Natural Silicon $Si^{nat}$

$92.2\% Si^{28}, 4.6\% Si^{29} 3.1\% Si^{32}$ is compared to the $Si^{28}$ enriched sample. The agreement is excellent.

### 4.3 Phonon Dynamics

The ability of the Monte Carlo approach to reproduce thermal transients has been tested on a thermal resistor with a length of 2 micrometers. The device has two lateral contacts set to temperatures $T_H = 310K$ and $T_L = 290K$, and the device is initialized at the temperature $T_L$. During the simulation, the transient effective temperature is extracted from the local energy density by inverting the energy vs. temperature relationship seen previously in equation (7).

The simulated evolution of the temperature can be estimated analytically by solving Fick’s law of diffusion. An approximate analytical solution of Fick’s law may be obtained by means of the Laplace transform:

$$T_F(x, t) = T_L + (T_H - T_L) \left[ \text{erfc} \left( \frac{x}{2\sqrt{tD}} \right) - \text{erfc} \left( \frac{2L - x}{2\sqrt{tD}} \right) + \text{erfc} \left( \frac{2L + x}{2\sqrt{tD}} \right) \right]$$  \hspace{1cm} (9)

where $T_F$ is the estimated temperature, $t$ the time, $D$ the diffusion coefficient, $L$ the length of the device, and $\text{erfc}$ is the complementary error function. The complementary error function $\text{erfc}$ is related to the error function by $\text{erfc}(z) = 1 - \text{erf}(z)$, and the error function itself is defined as:

$$\text{erf}(z) = \frac{2}{\sqrt{\pi}} \int_0^z e^{-t^2} dt$$  \hspace{1cm} (10)
Figure 4 shows both the theoretical temperature distribution (dashed lines) and the results of Monte Carlo simulation (solid lines) being averaged over 50 simulations (each one with different random number seed). The full band phonon Monte Carlo is seen to follow the estimation from Fick’s law at each time.

Figure 4: Comparison between the Theoretical Temperature Distribution (dashed lines), and the Monte Carlo Simulation (solid lines)

The simulations follow the analytical solution during the time transient.

4.4 Phonon-Electron Interaction

Within the CMC framework the scattering probabilities from all initial states to all the possible final states are pre-computed for a specific temperature and stored in a lookup scattering table. Whenever a scattering event occurs, the final state is chosen from the lookup table. This approach would require using a different scattering table to model different temperatures. A new structure for the scattering table and an enhanced scattering algorithm has been designed to model a wide range of temperatures using only one table while retaining the speed advantage of the CMC.

The new structure uses a two level approach. Each element of the first level array contains the information used in a typical Ensemble Monte-Carlo (EMC), e.g., the scattering rate, scattering mechanism, phonon mode involved, etc. Each element of the EMC array is linked to a second
array (CMC array), which contains all the possible final states and the scattering probability for that mechanism. The new algorithm requires two steps: first the new scattering table is pre-computed according to an expected maximum temperature (Tmax), second a rejection technique is used to adapt the scattering probability to the local temperature (Tloc).

Based on this scheme, the scattering is modeled as a five step process:

1. The scattering table is pre-computed for the maximum expected temperature Tmax.
2. A scattering mechanism is chosen from the EMC array
3. A final state is chosen from the related CA array
4. A rejection probability is computed according to the local runtime conditions
5. A stochastic process establishes whether the scattering occurs or is rejected.

Figure 5 shows the electron drift velocity versus electric field in gallium arsenide (GaAs) computed with the original CMC table structure (squares) at temperature T=150K, and the new structure with the rejection technique (solid lines) with Tmax=400K and Tloc=150K. The simulations overlap at low electric field, where the CMC grid is extremely fine. However, at high electric fields where the grid is coarser, the new structure produces a more accurate final state energy.

![Figure 5: Validation of the Rejection Algorithm by Simulation of the GaAs Electron Drift Velocity versus Electric Field](image)

The velocity computed using the rejection algorithm (red line) is identical to the one obtained with a fixed temperature table (dots). Further improvement is obtained by the implementation of the energy correction (dashed line)
Figure 6 shows the drift velocity versus electric field in Si at four different temperatures. The simulations employing the scattering table with the rejection technique (lines) are performed using the same pre-computed Tmax=350K rejection table and varying Tloc, while the reference simulations (dots) use a separate table for each temperature.

![Graph showing electron drift velocity versus electric field in Si at different temperatures.](image)

**Figure 6: Electron Drift Velocity versus Electric Field in Silicon at Different Temperature**  
*The lines are all computed using the same Tmax=350K scattering table and the rejection technique, whereas the velocities represented by dots are computed using fixed temperature tables.*

### 4.5 Beyond the Heat Equation: Energy Balance Equation Solver

The original research program proposed an efficient technique to reach thermal equilibrium within the device layout. This technique was based on a solution of the Heat Transport Equation (HTE) (flux-based) coupled self-consistently with the particle-based electron dynamics. Once the electro-thermal steady-state conditions were reached, the temperature map supplied by the HTE solver would have been replaced by a population of phonons and the particle-based phonon dynamics simulation engine would have been started in order to solve transients and non-equilibrium heat transport.

However, an approach based on solving the Energy Balance equation for phonons, in lieu of the Heat Transport Equation, was adopted. The Energy Balance approach is better suited to self-consistent coupling with the electron dynamics and allows a higher degree of accuracy by supplying a separate solution for each phonon mode (or group of modes). The approach developed here is based on the energy balance equation for each mode $\mu$, directly obtained from the phonon Boltzmann Transport Equation (BTE):

---

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\[
\frac{\partial W_\mu}{\partial t} = -\nabla \cdot F_\mu + \left. \frac{\partial W_\mu}{\partial t} \right|_{e-p} + \left. \frac{\partial W_\mu}{\partial t} \right|_{p-p}
\]  

(11)

where \( W_\mu(r, t) = \frac{1}{\Omega} \sum_k E_\mu(k) f_\mu(r, k, t) \) is the ensemble energy in the volume \( \Omega \) of the reciprocal space, \( F_\mu(r, t) = \frac{1}{\Omega} \sum_k v(k) E_\mu(k) f_\mu(r, k, t) \) is the energy flux, and the two partial derivatives of \( W_\mu \) in the LHS of the equation represent the rate of change of the ensemble energy density due to electron-phonon and phonon-phonon interaction, respectively.

Under steady-state conditions the time derivative on the LHS must be zero, and hence the heat flux is given by the sum of electron-phonon and phonon-phonon contributions, i.e.,

\[
\nabla \cdot \left( \kappa_\mu(T, r) \nabla T \right) = -\left( \left. \frac{\partial W_\mu}{\partial t} \right|_{e-p} + \left. \frac{\partial W_\mu}{\partial t} \right|_{p-p} \right) = -P_\mu
\]

(12)

where a Fourier Law approximation has been used for \( F_\mu(r) = -k_\mu(T, r) \nabla T \), and the subscript \( \mu \) denotes either the optical or acoustic phonon mode, respectively.

The CMC code includes efficient solvers for linear elliptical partial differential equations (PDEs), and thus it is desired to manipulate the energy balance equation in order to re-write it in the form of an elliptical PDE. The main issue with such manipulation is the dependency of \( k_\mu \) from the temperature and the position. We therefore assume that, with respect to the position, the thermal conductivity can be represented as a piece-wise function of the temperature, in other words, \( k_{\mu,C}(T) \) is a function of the temperature but is not changing with the position within each cell \( C \) of the finite differences grid. We therefore express this restricted position dependency with the index \( C \) rather than via a full functional dependence on the position vector \( r \).

With this in mind, we utilize the well-known Kirchhoff Transformation to define an “apparent” temperature \( \theta_{\mu,C}(T) \):

\[
\theta_{\mu,C}(T) = T_0 + \frac{1}{k_{\mu,C}(T_0)} \int_{T_0}^{T} k_{\mu,C}(\tau) d\tau
\]

(13)

where \( T_0 \) is a reference temperature, and \( k_{\mu,C}(T_0) \) the thermal conductivity of the phonon mode \( \mu \) at the reference temperature. This allows us to rewrite the energy balance equation as follows:

\[
\nabla^2 \theta_{\mu,C} = -\frac{P_\mu(r)}{k_{\mu,C}(T_0)}
\]

(14)

which is a linear Poisson equation for the temperature \( \theta_{\mu,C}(T) \). This final equation can be efficiently solved with the multi-grid solvers readily available in the code.

The temperature dependent thermal conductivity is available for many materials as a power law fit to experimental data [13]. Using these relationships, the Kirchhoff Transformation is computed within the code. This computation is performed for two groups of modes, the acoustic

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and the optical, and it is assumed that the temperature dependence affects only the acoustic modes. The temperature value along with its corresponding acoustic and optical Kirchhoff temperatures are tabulated for each material at the beginning of a simulation and used for the Kirchhoff and inverse Kirchhoff transformations throughout the simulation.

To validate the approach, we have solved the above equation with the forcing function set to zero, i.e., within the diffusive regime. In this case we expect the solution to simply be linear in between two boundary “thermal contacts” where the temperature is specified. To verify our approach, a thermal GaAs resistor has been simulated where a right contact is initially set at 310K, while a left contact and all intermediate cells are set to 300K. The temperature map is first transformed into a Kirchhoff apparent temperature, then the PDE is solved, and finally the inverse transformation is applied to return to the “standard” temperature variable. The resulting temperature map is linear as expected as seen in Figure 7.

Figure 7: Steady-State Temperature Distribution within a Device with the Left Contact set at 300K and the Right one at 310K

The solution is linear in the region within the contacts as expected in the absence of a forcing function.
Next, a manner to calculate the forcing function for the energy balance equation must be implemented. The forcing function for a phonon mode \( \mu \) can be separated into individual contributions due to electron-phonon (e-p) scattering and to phonon-phonon (p-p) scattering, representing the rate of change of the phonon energy density for the particular mode, respectively. The forcing function \( P_\mu \) is then written as:

\[
- P_\mu = - \left( \frac{\partial W_\mu}{\partial t} \bigg|_{e-p} + \frac{\partial W_\mu}{\partial t} \bigg|_{p-p} \right)
\]  

(15)

The electron-phonon term in the forcing function can be computed by tracking the energy exchanged between phonons and electrons during scattering, while the phonon-phonon term is expressed using a relaxation time approximation [14]:

\[
\frac{\partial W_\mu}{\partial t} \bigg|_{op-ac} = -C_{op} \frac{T_{op} - T_{ac}}{\tau_{op-ac}},
\]  

(16)

\[
\frac{\partial W_\mu}{\partial t} \bigg|_{ac-op} = -C_{ac} \frac{T_{ac} - T_{op}}{\tau_{ac-op}},
\]  

(17)

where \( C_\mu \) is the volumetric heat capacity \( \left( \frac{J}{m^3 K} \right) \), \( T_\mu \) the temperature of mode \( \mu \), and \( \tau_{i-j} \) the energy relaxation time between mode \( i \) and \( j \). In the above equations, the subscripts \( ac \) and \( op \) denote the acoustic and optical modes, respectively.

### 4.6 Boundary Conditions

It was initially assumed that a simple Dirichlet boundary condition [15] [16] on a thermal contact would be sufficient for our simulation purposes. In the Dirichlet boundary condition, the temperature value is simply set on the boundaries where specified and not allowed to change from this prescribed value, i.e., \( T_b = c \)

However we observed that, both in the Monte Carlo simulations as well as the commercial simulator Synopsys [17], that this condition seemed to cool the system too much with peak temperatures in a biased device rising only a few Kelvin above room temperature. Other boundary conditions explored are those of the inhomogeneous Neumann [15] [16] and the Robin [15] [16] boundary conditions. It should be noted that the Robin boundary condition is also called a generalized Neumann condition within the Matlab [18] software PDE package. In addition, the Dirichlet boundary condition is sometimes called a boundary condition of the first kind, the Neumann boundary condition a boundary condition of the second kind, and the Robin boundary condition a boundary condition of the third kind [16].

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The Neumann condition is that of prescribing the heat flux that leaves the specified boundary, and is expressed as:

$$\frac{\partial \nu}{\partial \mathbf{n}} = f$$  \hfill (18)

while the related homogenous Neumann condition, for instance that of an insulating surface, is expressed as:

$$\frac{\partial T}{\partial n} = 0$$  \hfill (19)

The problem with this condition is that it does not produce a unique solution, but rather a family of solutions differing by an additive constant of integration. Hence, to effectively use this condition the temperature must be prescribed at some location within the simulation domain $R$, i.e., there must be at least one point within the device in question at which we have a Dirichlet boundary condition. In addition the elliptic PDE with a Neumann boundary condition:

$$-\nabla^2 v = F \text{ in } R$$  \hfill (20)

$$\frac{\partial v}{\partial n} = g \text{ on } \partial R$$  \hfill (21)

is also subject to the following constraint [15]:

$$-\int_R F \, d(x, y) = \int_{\partial R} g \, dS.$$  \hfill (22)

The Robin boundary condition is expressed as

$$-k \frac{\partial \nu}{\partial \mathbf{n}} = h(T - T_a)$$  \hfill (23)

where $T$ is the interior temperature at the boundary and $T_a$ is the temperature of the outside environment or, in our case, the temperature of the heat sink of the device. The heat transfer coefficient, $h$, has units $W/m^2K$ and can be thought of as a thermal surface conductance, or the inverse of a thermal surface resistance. The role of the heat transport coefficient is to scale the outwardly directed flux based on how well heat should be dissipated away from the boundary. The Robin condition yields a unique solution as long as $\frac{k}{h} > 0$ [15].

The Dirichlet and Robin boundary conditions have been implemented and tested using a simple case with an analytic solution. In this test, a uniform piece of GaAs is used with either the Dirichlet or Robin condition set on the bottom plane of the material and a homogenous Neumann condition set on all other surfaces (meaning that zero heat flux leaves those surfaces).
The governing differential equation is a Poisson Equation of the form (in 1-D)

\[
\frac{d^2T}{dy^2} = -\frac{q}{\kappa},
\]

which has a simple analytical solution of

\[
T = -\frac{q}{2\kappa} y^2 + C_1 y + C_2
\]  \hspace{1cm} (25)

The constant \( C_1 = \frac{q^* L}{k} \), where \( L \) is the length of the heat-generating material, for both the Dirichlet and Robin boundary conditions. For the Dirichlet condition, the constant \( C_2 \) is given by the temperature specified at the boundary, \( C_2 = T_b \), while for the Robin condition it is found that \( C_2 = T_a + \frac{q^* L}{h} \).

Various values for the uniform heat generation rate \( q \) have been used in the Dirichlet case, while different values of the transfer coefficient \( h \) were used in the Robin case. Simulations were then carried out considering the cases of a temperature independent thermal conductivity (as for the analytic solution) and that of a temperature dependent thermal conductivity.

The results of these tests are shown in Figure 8 and Figure 9, respectively. The numerical results for the temperature independent case are seen to overlap the expected analytic solutions. Using a temperature-dependent thermal conductivity in the simulation results in a rise in the temperature distribution of approximately 2\% in the case of a Robin condition with \( h = 2e5 \frac{W}{m^2} \), 0.08\% with \( h = 1e7 \frac{W}{m^2} \), 0.8\% in the case of a Dirichlet condition with \( q = 5.34e16 \frac{W}{m^3} \), and 0.0006 \% with \( q = 5.34e13 \frac{W}{m^3} \). This rise in temperature is expected due to a decrease in thermal conductivity as the temperature increases. The difference between the constant and temperature dependent-thermal conductivity temperature distributions also increases as the temperatures become higher, e.g., a roughly 2\% difference for the Robin condition with \( h = 2e5 \frac{W}{m^2} \) with \( h = 2e5 \frac{W}{m^2} \) where the peak temperature for the constant case is 350.75K but only a 0.08\% change with \( h = 1e7 \frac{W}{m^2} \) where the peak temperature is 301.27K. Due to the weak non-linearity in the \( k \) vs \( T \) relationships, the accuracy of the simulated result in the constant \( k \) case, and the behavior of the changes in the temperature distributions of the dependent \( k \) simulations we conclude that the use of our Kirchhoff transformation approach is justified.

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Figure 8: Results for the Dirichlet Boundary Condition
The constant thermal conductivity simulation (red) is seen to overlap the analytic solution (blue), while the simulation with a temperature-dependent thermal conductivity reaches higher peak temperatures as expected due the inverse relationship between thermal conductivity and temperature.

Figure 9: Results for the Robin Boundary Condition
The constant thermal conductivity simulation (red) overlaps the analytic solution (blue), while the simulation using a temperature-dependent thermal conductivity reaches higher peak temperatures due to the inverse relationship between thermal conductivity and temperature.

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5.0 Conclusions

This project presented a novel modeling approach for the development of a technology based on the microscopic thermal management of solid-state devices. New techniques that model the phonon dynamics within a full spectrum framework have been presented along the comparison between simulated and experimental results. A rejection technique has been implemented to adapt the electron scattering rates to the local conditions, and the thermal conductivity has been correctly reproduced using phonon-phonon and isotope scattering within a particle based approach.

In addition, an energy balance approach derived from the phonon BTE was successfully manipulated into the form of an elliptic PDE through the implementation of the Kirchhoff Transformation, and agreement with analytical solutions was demonstrated. Boundary conditions relevant to thermal simulations were also explored resulting in the implementation of the Robin boundary condition.
6.0 References


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