Testing Coordinate Frame Transformations
NOVAS vs SOFA

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The Standards of Fundamental Astronomy (IAU 2009a; SOFA)\(^1\) library is the official collection of approved software for positional astronomy, operating under the auspices of International Astronomical Union (IAU) Division 1 (Fundamental Astronomy). Both Fortran and C libraries are available. An international SOFA Reviewing Board manages the collection.

Generally, the Naval Observatory Vector Astrometry Software (Kaplan et al. 2009; NOVAS)\(^2\) is independent of SOFA although both software libraries include code that is similar to two International Earth Rotation and Reference Systems Service (IERS) Fortran modules.\(^3\) NU2000A and iau2000a (Fortran and C, respectively), which evaluate the full 1,365-term IAU 2000A nutation series in NOVAS 3.0, are based on the IERS subroutine NU2000A. EECT2000 and ee_ct, which evaluate the “complementary terms” in the equation of the equinoxes, are based on IERS function EECT2000. The corresponding modules in SOFA are iau_NUT00A, iauNut00a, iau_EECT00, and iauEect00.

The document SOFA Tools for Earth Attitude (IAU 2009b), also known as the “SOFA Cookbook,” contains several Fortran examples of the transformation between terrestrial and celestial coordinate systems. This technical note examines how one of those examples plays out in both NOVAS and SOFA.

Goal

These tests were designed to compare the transformation from the Geocentric Celestial Reference System (GCRS) to International Terrestrial Reference System (ITRS) using the IAU 2000A/2006 models for precession and nutation. For the Fortran test, we compared the Celestial Intermediate Origin (CIO) based method in NOVAS_F3.0g\(^4\) at full accuracy\(^5\) with the “IAU 2006/2000A, CIO based, using classical angles” example in the SOFA Cookbook. For the C test, we used SOFA_Test.c, attached to this report. The goal was to verify that the NOVAS libraries, which are (mostly) independent of SOFA, produced results that agree with

\(^1\) http://www.iausoфа.org/index.html
\(^2\) http://www.usno.navy.mil/USNO/astronomical-applications/software-products/novas
\(^3\) http://tai.bipm.org/iers/conv2003/conv2003_c5.html
\(^4\) beta version “g” of NOVAS F3.0
\(^5\) CALL ciotio/CALL hiacc for mode = 0
Testing Coordinate Frame Transformations NOVAS vs SOFA

The Standards of Fundamental Astronomy (SOFA) library is the official collection of approved software for positional astronomy, operating under the auspices of International Astronomical Union (IAU) Division 1 (Fundamental Astronomy). Both Fortran and C libraries are available. An international SOFA Reviewing Board manages the collection. Generally, the Naval Observatory Vector Astrometry Software (NOVAS) is independent of SOFA although both software libraries include code that is similar to two International Earth Rotation and Reference System Service (IERS) Fortran modules, NU2000A and iau2000a (Fortran and C, respectively), which evaluate the full 1,365-term IAU 2000A nutation series in NOVAS 3.0, are based on the IERS subroutine NU2000A. EECT2000 and ee_ct, which evaluate the "complementary terms" in the equation of the equinoxes, are based on IERS function EECT2000. The corresponding modules in SOFA are iau_NUT00A, iauNut00a, iau_EECT00, and iauEect00. The document SOFA Tools for Earth Attitude, also known as the "SOFA Cookbook", contains several Fortran examples of the transformation between terrestrial and celestial coordinate systems. This technical note examines how one of those examples plays out in both NOVAS and SOFA.
their SOFA counterparts at a level that is at least an order of magnitude better than the best observational results.

Procedure for Fortran Tests

The programs used for the Fortran tests were TERCEL-TEST.f and SOFA_TEST.f; copies of which are in Addendum I and Addendum II, respectively. The input parameters, which were taken from the SOFA Cookbook, were the following:

- **Universal Time**: \( UT1 = 2400000.5 + 54195.499991658 \) days (Julian date)
  The UT1 value is divided into two parts, i.e., two separate arguments, because of the large number of significant digits needed for precise results. The best agreement between NOVAS and SOFA was obtained when UT1 was split in the exactly the same place; splitting the date differently produced differences of about 3 microarcseconds (\( \mu \)as).

- **Difference between Terrestrial Time (TT) and UT1**: \( \Delta T = 65.25607389 \) s
  SOFA does not use \( \Delta T \); this value is the difference between the TT and UT1 Julian dates in the SOFA example, expressed in seconds.

- **Polar coordinates**: \( XP = 0.0349282, YP = 0.4833163 \) arcsec

- **Celestial Intermediate Pole (CIP) offsets**: \( DX = 0.1725, DY = -0.265 \) arcsec

  The SOFA subroutine \( iau_NUT06A \) includes small corrections to the nutation series arising from the P03 precession (Capitaine, Wallace, & Chapront 2003; hereafter P03) that are not used in the NOVAS calculations. The corrections amount to only a few microarcseconds for current dates.

NOVAS does not directly produce an overall GCRS-to-ITRS rotation matrix as SOFA does. The NOVAS rotation matrix was constructed simply by passing the three vectors, \((1,0,0)\), \((0,1,0)\), and \((0,0,1)\), in succession through subroutine TERCEL.

A series of tests were done, with and without corrections for polar motion, precession and nutation, and the P03 correction in SOFA, and the resulting rotation matrices were compared. The computations were executed on a 32-bit Mac system running the Leopard (OS X 10.5) operating system.

Results of Fortran Tests

Table 1 shows that the latest Fortran releases of NOVAS and SOFA agree at the sub-microarcsecond level in the transformation between the celestial and terrestrial reference systems when the same Earth orientation parameters and conventions are used. In this case, including the P03 corrections in the SOFA nutation adds a discrepancy on the order of 1.4 \( \mu \)as. Inclusion of the CIP offsets and polar motion does not significantly add to the differences in the two formulations, as long as the parameters used are identical in the two cases. Use of the external CIO_RA file in the NOVAS calculation adds about 0.05 \( \mu \)as to the difference for the above case, while using equinox mode for the NOVAS computations does not have a significant effect on the results.
Table 1. Comparison of NOVAS F3.0 and SOFA Fortran

<table>
<thead>
<tr>
<th>Polar Motion</th>
<th>CIP Offsets</th>
<th>P03 Terms</th>
<th>External CIO_RA</th>
<th>EQINOX mode</th>
<th>Difference (µas)</th>
</tr>
</thead>
<tbody>
<tr>
<td>No</td>
<td>No</td>
<td>No</td>
<td>No</td>
<td>No</td>
<td>0.25814</td>
</tr>
<tr>
<td>No</td>
<td>No</td>
<td>Yes</td>
<td>No</td>
<td>No</td>
<td>1.6752</td>
</tr>
<tr>
<td>Yes</td>
<td>Yes</td>
<td>Yes</td>
<td>No</td>
<td>No</td>
<td>1.6728</td>
</tr>
<tr>
<td>Yes</td>
<td>Yes</td>
<td>No</td>
<td>No</td>
<td>No</td>
<td>1.6735</td>
</tr>
<tr>
<td>Yes</td>
<td>Yes</td>
<td>No</td>
<td>Yes</td>
<td>No</td>
<td>0.28679</td>
</tr>
<tr>
<td>Yes</td>
<td>Yes</td>
<td>No</td>
<td>No</td>
<td>Yes</td>
<td>0.34369</td>
</tr>
<tr>
<td>Yes</td>
<td>Yes</td>
<td>No</td>
<td>No</td>
<td>Yes</td>
<td>0.28644</td>
</tr>
</tbody>
</table>

The results presented in Table 1 were obtained by computing the GCRS to ITRS transformations for the single time discussed in the SOFA Cookbook. Therefore, the values should be typical. Comparisons of future releases should compare transformations for multiple times to estimate the range of possible differences.

Procedure for C Tests

The program used for the C tests was SOFA_Test.c; a copy of which is in Addendum III. SOFA_Test.c is basically a line-for-line transliteration of the Fortran SOFA-TEST.f that uses the SOFA C functions. At the end of the file, a C function terceltest, based on the Fortran program terceltest, is included. C function terceltest was run separately from the bulk of SOFA_Test.c using the following input parameters:

- **Universal Time:** $\text{UT1} = 2400000.5 + 54195.4999991658$ days (Julian date)
  The UT1 value is divided into two parts, i.e., two separate arguments, because of the large number of significant digits needed for precise results. The best agreement between NOVAS and SOFA was obtained when UT1 was split in the exactly the same place.

- **Difference between TT and UT1:** $\Delta T = 65.25607389$ s
  SOFA does not use $\Delta T$; this value is the difference between the TT and UT1 Julian dates in the SOFA example, expressed in seconds.

- **Polar coordinates:** $X_P = 0.0349282, Y_P = 0.4833163$ arcsec

- **CIP offsets:** $D_X = 0.1725, D_Y = -0.265$ arcsec

The SOFA function iauNut06a includes small corrections to the nutation series arising from the P03 precession that are not used in the NOVAS calculations. The corrections amount to only a few microarcseconds for current dates.

NOVAS does not directly produce an overall GCRS-to-ITRS rotation matrix as SOFA does. The NOVAS rotation matrix was constructed simply by passing the three vectors, $(1,0,0)$, $(0,1,0)$, and $(0,0,1)$, in succession through function ter2cel.
A series of tests were done, with and without corrections for polar motion, precession and nutation, and the P03 correction in SOFA. The output of the C version of terceltest and SOFA Test were compared with output from the corresponding Fortran programs. As with the Fortran tests, the C computations were executed on a 32-bit Mac system running the Leopard (OS X 10.5) operating system.

Results of the C Tests

The outputs of the SOFA C tests were identical with those produced by the Fortran versions. Addendum III includes a sample output from the C tests.

References


IAU. 2009a, Standards of Fundamental Astronomy, (SOFA) http://www.iausofa.org/


IERS. 2003, Conventions 2003: Chapter 5 Transformation Between the Celestial and Terrestrial Systems (Frankfurt, Germany: BKG)

http://tai.bipm.org/iers/conv2003/conv2003_e5.html

\(^6\) http://www.usno navy.mil/USNO/astronomical-applications/publications/circ-180
Addendum I: NOVAS Fortran Test Code

The NOVAS Fortran code for the comparison of NOVAS and SOFA is contained in PROGRAM terceltest below. This program reads the contents of tercel-test-input.dat, which is also listed below. The output of terceltest may be directed to a file, such as tercel_matrix, which can then be read by SOFA-TEST.f.

Program File: TERCEL-TEST.f

PROGRAM terceltest

C
C
C---PURPOSE: Transform vectors from ITRS to GCRS for comparing NOVAS with SOFA
C
C---REFERENCES: None
C
C---INPUT
C   ARGUMENTS: num spacer read from file
C                (INTEGER)
C   tjdh TDB or TT Julian Date, higher part, read from file
C                (DOUBLE PRECISION)
C   tjd1 TDB or TT Julian Date, lower part, read from file
C                (DOUBLE PRECISION)
C   xp conventionally-defined x coordinate of CIP
C   with respect to ITRS pole, arcseconds, read from file
C                (DOUBLE PRECISION)
C   yp conventionally-defined y coordinate of CIP
C   with respect to ITRS pole, arcseconds, read from file
C                (DOUBLE PRECISION)
C   vec1 position vector, geocentric equatorial rectangular
C   coordinates, referred to ITRS axes, read from file
C                (DOUBLE PRECISION)
C
C---OUTPUT
C   ARGUMENTS: vec2 position vector, geocentric equatorial rectangular
C   coordinates, referred to GCRS axes
C                (DOUBLE PRECISION)
C
C---COMMON
C   BLOCKS: None
C
C---ROUTINE
C   CALLED: SUBROUTINE setdt    (NOVAS)
C   SUBROUTINE celpol   (NOVAS)
C   SUBROUTINE ciotio   (NOVAS)
C   SUBROUTINE eqinox   (NOVAS)
C   SUBROUTINE hiacc    (NOVAS)
C   SUBROUTINE tercel   (NOVAS)
C
C---COMPILING: compile with NOVAS, solsys2, jpljubs
C   have JPLEPH in working directory
have CIO_RA.TXT in working directory for tests w/ external CIO_RA file

C---VER./DATE/
C PROGRAMMER: v1.0/02-09/AM (USNO/AA) initial version
C v1.5/09-10/JLB (USNO/AA) additional documentation
C
C---NOTES       1. Input arguments read from tercel-test-input.dat.
C               2. For best agreement with SOFA, split the date at
C                  the same point, i.e. tjdh = 2400000.5
C               3. Send output to file, tercel_matrix, to be read by
C                  SOFA_TEST.f, which does comparison.
C               4. Initial test variations
C                  A. no polar motion, no CIP offsets, no P03
C                     corrections (SOFA), no CIO_RA, cio mode
C                  B. no polar motion, no CIP offsets, P03
C                     corrections (SOFA), no CIO_RA, cio mode
C                  C. no polar motion, CIP offsets, P03
C                     corrections (SOFA), no CIO_RA, cio mode
C                  D. polar motion, CIP offsets, P03
C                     corrections (SOFA), no CIO_RA, cio mode
C                  E. polar motion, CIP offsets, no P03
C                     corrections (SOFA), no CIO_RA, cio mode
C                  F. polar motion, CIP offsets, no P03
C                     corrections (SOFA), CIO_RA, cio mode
C                  G. polar motion, CIP offsets, no P03
C                     corrections (SOFA), no CIO_RA, equinox mode
C---------------------------------------------------------------------
DOUBLE PRECISION tjdh, tjdl, xp, yp, delt, vec1(3),  
     vec2(3), tjd, dx, dy
INTEGER num, mode

DATA delt /65.25607389d0/, num / 0 /
CALL setdt (delt)

C-----dx, dy are celestial pole offsets to be used
C-----use 1st pair to invoke correction
C-----use 2nd pair to ignore correction
C
dx = +0.1750d0
dy = -0.2259d0
C dy = 0d0
C dx = 0d0

C-----Open the input file of Julian dates,CIO coords,ITRS vector

OPEN (UNIT = 15, FILE = 'tercel-test-input.dat',  
     STATUS = 'OLD', ACCESS = 'SEQUENTIAL')

10 READ (15,*, END = 20) num, tjdh, tjdl, xp, yp, vec1

C-----Set transformation method, accuracy level, and UT1-UTC.
C-----CIO-based method (ciotio) is used for most tests
C-----Equinox-base method (eqinox) is used in one case
tjd = tjdh + tjdl

CALL celpol (tjd, 2, dx, dy)

CALL ciotio
CALL eqinox
CALL hiacc

C-----Rotate vec1 from ITRS to GCRS = vec2

CALL TERCEL(tjdh, tjdl, xp, yp, vec1, vec2)
WRITE (*, 101) num, vec2
GO TO 10

101 FORMAT (i1, 3(4x, f20.17))

STOP
END

Input Data File: tercel-test-input.dat
1 2400000.5 54195.4999916581146 +0.0349282d0 +0.4833163d0 1d0 0d0 0d0
1 2400000.5 54195.4999916581146 +0.0349282d0 +0.4833163d0 0d0 1d0 0d0
1 2400000.5 54195.49999916581146 +0.0349282d0 +0.4833163d0 0d0 0d0 1d0
Addendum II: SOFA Fortran Test Code

The SOFA Fortran code for the comparison of NOVAS and SOFA is contained in SOFA_TEST.f, which is based on Example 5.5 “IAU 2006/2000A, CIO based, using classical angles” in the SOFA Cookbook; a copy of the program file follows. The program reads two input data files: tercel_matrix and sofa_matrix, which are listed below. At the end of this section, a sample of its output is provided.

**Program File: SOFA-TEST.f**

```fortran
PROGRAM SOFA_Test

* * *---PURPOSE: Compare performance of NOVAS F3.0 & SOFA Fortran libraries * * *---REFERENCES: IAU. 2009, SOFA Tools for Earth Attitude, Software * version 4, Document revision 1.1 (SOFA Cookbook) * http://www.iausofa.org/sofa pn.pdf * * *---INPUT * ARGUMENTS: TERMAT  =  3x3 matrix read from tercel_matrix * (DOUBLE PRECISION) * SOFMAT  =  3x3 matrix read from sofa_matrix * (DOUBLE PRECISION) * *---OUTPUT * ARGUMENTS: NONE * * *---COMMON * BLOCKS: NONE * * *---ROUTINES * CALLED: DOUBLE PRECISION FUNCTION iau_ANP (SOFA) * SUBROUTINE iau_BPN2XY (SOFA) * SUBROUTINE iau_C2IXYS (SOFA) * SUBROUTINE iau_C2IXYS (SOFA) * SUBROUTINE iau_CAL2JD (SOFA) * SUBROUTINE iau_CP (SOFA) * SUBROUTINE iau_CR (SOFA) * SUBROUTINE iau_D2TF (SOFA) * SUBROUTINE iau_DAT (SOFA) * DOUBLE PRECISION FUNCTION iau_ERA00 (SOFA) * DOUBLE PRECISION FUNCTION iau_PAD03 (SOFA) * DOUBLE PRECISION FUNCTION iau_FAE03 (SOFA) * DOUBLE PRECISION FUNCTION iau_FAF03 (SOFA) * DOUBLE PRECISION FUNCTION iau_FAJU03 (SOFA) * DOUBLE PRECISION FUNCTION iau_FAL03 (SOFA) * DOUBLE PRECISION FUNCTION iau_FALP03 (SOFA) * DOUBLE PRECISION FUNCTION iau_FAMA03 (SOFA) * DOUBLE PRECISION FUNCTION iau_FAME03 (SOFA) * DOUBLE PRECISION FUNCTION iau_FAOM03 (SOFA) * DOUBLE PRECISION FUNCTION iau_FAPA03 (SOFA) * DOUBLE PRECISION FUNCTION iau_FASA03 (SOFA)
```
* DOUBLE PRECISION FUNCTION iau_FAUR03 (SOFA)
* DOUBLE PRECISION FUNCTION iau_FAVE03 (SOFA)
* SUBROUTINE iau_FW2M (SOFA)
* SUBROUTINE iau_IR (SOFA)
* SUBROUTINE iau_NUT00A (SOFA)
* SUBROUTINE iau_NUT06A (SOFA)
* DOUBLE PRECISION FUNCTION iau_OBL06 (SOFA)
* SUBROUTINE iau_PFW06 (SOFA)
* SUBROUTINE iau_PM (SOFA)
* SUBROUTINE iau_PNM06A (SOFA)
* SUBROUTINE iau_POM00 (SOFA)
* SUBROUTINE iau_RM2V (SOFA)
* SUBROUTINE iau_RX (SOFA)
* SUBROUTINE iau_RXR (SOFA)
* SUBROUTINE iau_RY (SOFA)
* SUBROUTINE iau_RZ (SOFA)
* DOUBLE PRECISION FUNCTION iau_S06 (SOFA)
* DOUBLE PRECISION FUNCTION iau_SP00 (SOFA)
* SUBROUTINE iau_TR (SOFA)
* SUBROUTINE iau_XY06A (SOFA)
* SUBROUTINE iau_ZR (SOFA)
* SUBROUTINE REPMAT (internal)
* DOUBLE PRECISION FUNCTION DROT (internal)
* SUBROUTINE TRANSPOSE (internal)
* *---COMPILING: SOFA_Test.f must be compiled with appropriate SOFA
* modules. SOFA modules are available as discrete
* files. Downloading the entire library and
* concatenating it into a single file may be easiest.
* *---VER./DATE/
* PROGRAMMER: v1.0/02-09/AM (USNO/AA) initial version
* v1.1/09-10/JLB (USNO/AA) additional documentation
* *---NOTES 1. Input arguments read from tercel_matrix, sofa_matrix
* 2. SOFA_Test does not call all of the routines listed
* above directly; some are called by routines called by
* SOFA_Test. A fuller than usual list is given here to
* ensure the user has all the necessary SOFA files.
* *---------------------------------------------

* Preliminaries from SOFA Cookbook, section 5.1

* SOFA examples

IMPLICIT NONE

* Arcseconds to radians
DOUBLE PRECISION AS2R
PARAMETER ( AS2R = 4.848136811095359935899141D-6 )

* 2Pi
DOUBLE PRECISION D2PI
PARAMETER ( D2PI = 6.283185307179586476925287D0 )
* PM, R1(3,3) through R5(3,3), TERMAT (3,3), TERTRANS(3,3),
* SOFMAT, SOFTRANS, DROT, DJMJDATE, UTHI, UTLO
* are not from Cookbook

CHARACTER PM
INTEGER IY, IM, ID, IH, MIN, J
INTEGER IHMSF(4)
DOUBLE PRECISION SEC, XP, XP, DUT1,
: DDP80, DDE80, DX00, DY00, DX06, DY06,
: DJMJD0, DATE, TIME, UTC, DAT,
: TAI, TT, TUT, UT1, RP(3,3), DP80, DE80,
: DPSI, DEPS, EPSA, RN(3,3), RNPB(3,3),
: EE, GST, RC2TI(3,3), RPOM(3,3),
: RC2IT(3,3), X, Y, S,
: RC2I(3,3), ERA, DP00, DE00, RB(3,3),
: DPB(3,3), V1(3), V2(3), DDP00, DDE00
DOUBLE PRECISION R1(3,3), R2(3,3), R3(3,3), R4(3,3), R5(3,3)
DOUBLE PRECISION iau_OBL80, iau_EQEQ94, iau_ANP, iau_GMST82,
: iau_ERA00, iau_SP00, iau_EE00, iau_GMST00,
: iau_S06,TERMAT(3,3), TERTRANS(3,3),
: SOFMAT(3,3), SOFTRANS(3,3)
DOUBLE PRECISION DROT, DJMJDATE, UTHI, UTLO

* Open and read file containing output from tercel.f
C.....
OPEN (UNIT = 16, FILE = 'tercel_matrix', STATUS = 'OLD',
. ACCESS = 'SEQUENTIAL')
READ (16,90) TERMAT
90 FORMAT ( 1x, 3e24.12 )
WRITE (*,'( 3(f20.15) )') TERMAT

* Open and read file containing sofa_matrix
C
OPEN (UNIT = 17, FILE = 'sofa_matrix', STATUS = 'OLD',
. ACCESS = 'SEQUENTIAL')
READ (17,95) SOFMAT
95 FORMAT ( 1x, 3f24.17 )
C WRITE (*,'( 3(f20.15) )') SOFMAT
C.....

* Initialize variables using values from SOFA Cookbook Preliminaries

* UTC.
IY = 2007
IM = 4
ID = 5
IH = 12
MIN = 0
SEC = 0D0
WRITE ( *, '(1X,''date'','I6.4,2('''',I2.2))'' ) IY, IM, ID
WRITE ( *, '(1X,''time''I4,I3,F5.1,'' UTC'','')'' ) IH, MIN, SEC

* Polar motion (arcsec->radians).
XP = 0.0349282D0 * AS2R
YP = 0.4833163D0 * AS2R  
WRITE ( *, '=,'/(1X,'''x_p '='',SP,F13.9,''' arcsec''')' ) XP/AS2R  
WRITE ( *, '=,'/(1X,'''y_p '='',SP,F13.9,''' arcsec''')' ) YP/AS2R  

* UT1-UTC (s).  
DUT1 = -0.072073685D0  
WRITE ( *, '=,'/(1X,'''UT1-UTC '='',SP,F16.12,''' s''')' ) DUT1  

DDP80 = -55.0655D0 * AS2R/1000D0  
DDE80 = -6.3580D0 * AS2R/1000D0  
WRITE ( *, '=,'/(1X,''''dDpsi (1980) '='',SP,F13.9,''' arcsec''')' ) : DDP80 / AS2R  
WRITE ( *, '=,'/(1X,''''dDepsi (1980) '='',SP,F13.9,''' arcsec''')' ) : DDE80 / AS2R  

* CIP offsets wrt IAU 2000A (mas->radians).  
DX00 = 0.1725D0 * AS2R/1000D0  
DY00 = -0.2650D0 * AS2R/1000D0  
WRITE ( *, '=,'/(1X,''''dX (2000) '='',SP,F13.9,''' arcsec''')' ) : DX00 / AS2R  
WRITE ( *, '=,'/(1X,''''dY (2000) '='',SP,F13.9,''' arcsec''')' ) : DY00 / AS2R  

* CIP offsets wrt IAU 2006/2000A (mas->radians).  
* First set from SOFA Cookbook, 2nd set to ignore correction  
DX06 = 0.1750D0 * AS2R/1000D0  
DY06 = -0.2259D0 * AS2R/1000D0  
C DX06 = 0d0  
C DY06 = 0d0 * AS2R/1000D0  
WRITE ( *, '=,'/(1X,''''dX (2006) '='',SP,F13.9,''' arcsec''')' ) : DX06 / AS2R  
WRITE ( *, '=,'/(1X,''''dY (2006) '='',SP,F13.9,''' arcsec''')' ) : DY06 / AS2R  

* TT (MJD).  
CALL iau_CAL2JD ( IY, IM, ID, DJMJD0, DATE, J )  
TIME = ( 600D0*(600D0*DBLE(IH) + DBLE(MIN)) + SEC ) / 86400D0  
UTC = DATE + TIME  
CALL iau_DAT ( IY, IM, ID, TIME, DAT, J )  
TAI = UTC + DAT/86400D0  
TT = TAI + 32.184D0/86400D0  
WRITE ( *, '=,'/(1X,''''TT = 2400000.5 +''',F22.15 ,')' ) TT  

* UT1.  
TUT = TIME + DUT1/86400D0  
UT1 = DATE + TUT  
WRITE ( *, '=,'/(1X,''''UT1 = 2400000.5 +''',F22.15 ,')' ) UT1  

* Following SOFA Cookbook, IAU 2006/2000A, CIO based,  
* classical angles, section 5.5 with additional intermediate output  

* ==============
* IAU 2006/2000A, CIO-based
* =========================

WRITE ( *, '(/1X,''   =========================''' //
:            '/1X,''4) IAU 2006/2000A, CIO-based''' //
:            '/1X,''   ========================='')' )

* CIP and CIO, IAU 2006/2000A.
C-----Report inputs to CIP calculation

CALL iau_XYS06A ( DJMJD0, TT, X, Y, S )

* Add CIP corrections.
  X = X + DX06
  Y = Y + DY06

* Report CIP and s.
  WRITE ( *, '(/1X,''X ='',SP,F22.15)' ) X/AS2R
  WRITE ( *, '( 1X,''Y ='',SP,F22.15)' ) Y/AS2R
  WRITE ( *, '( 1X,''s ='',SP,F13.9,'' arcsec'')' ) S/AS2R

* GCRS to CIRS matrix.
  CALL iau_C2IXYS ( X, Y, S, RC2I )

* Report.
  CALL REPMAT ( 'NPB matrix, CIO-based (RC2I)', RC2I )

* Earth rotation angle.
C
C----- Set TUT to the value used in Tercel

WRITE (*,*) 'DJMJD0, DATE, TUT: ', DJMJD0, DATE, TUT
C     DJMJD0 = 2400000.5d0
C     TUT = 54195.4999991658d0
C     UTHI = 2454195.0d0
C     UTLO = 0.99999916581146d0
C
ERA = iau_ERA00 ( DJMJD0, DATE+TUT )

WRITE ( *, '(1X)' ) ERA
C     ERA = iau_ERA00 ( UTHI, UTLO )
C     ERA = iau_ERA00 ( DJMJD0+DATE, TUT )

WRITE ( *, '(1X)' ) ERA/D2PI
C     ERA = iau_ERA00 ( DJMJD0, DATE+TUT )
C     ERA = iau_ERA00 ( UTHI, UTLO )
C     ERA = iau_ERA00 ( DJMJD0+DATE, TUT )

WRITE ( *, '(1X)' ) ERA
WRITE ( *, '(1X,''ERA ='',F19.16,'' rad'' )' ) ERA
WRITE ( *, '(1X,''ERA ='',F16.12,'' deg'' )' ) ERA*360D0/D2PI
CALL iau_D2TF ( 9, ERA/D2PI, PM, IHMSF )
WRITE ( *, '(1X,''PM ='',3I3.2,''.'',I9.9 )' ) IHMSF

* Form celestial-terrestrial matrix (no polar motion yet).
  CALL iau_CR ( RC2I, RC2TI )

CALL iau_RZ ( ERA, RC2TI )

* Report.
CALL REPMAT ( 'celestial to terrestrial matrix (no polar motion)',
              RC2TI )
CALL iau_CR (RC2TI,R3)

CALL iau_POM00 ( XP, YP, iau_SP00(DJMJD0,TT), RPOM )

* Form celestial-terrestrial matrix (including polar motion).
CALL iau_RXR ( RPOM, RC2TI, RC2IT )

* Compare results
C Transpose TERMAT
   CALL TRANSPOSE ( TERMAT, TERTRANS )
   CALL TRANSPOSE ( SOFMAT, SOFTRANS )

* Report.
   CALL REPMAT ( 'celestial to terrestrial matrix', RC2IT )
   CALL REPMAT ( 'tercel_transposed matrix', TERTRANS )

C   WRITE ( *, '(1X)' )
C   WRITE (*,'( 3(f20.15) )') TERMAT

* Copy for later comparison.
   CALL iau_CR ( RC2IT, R4 )

* Compare result to TERCEL_Transposed matrix
   WRITE ( *, '(1X)' )
   WRITE (*,'( 1x,''w/ pm result vs tercel ='',e20.10,'' uas'' )')
   . DROT ( R4, TERTRANS ) *1D6 / AS2R

* Compare SOFA w/p03 and SOFA w/o P03
   WRITE ( *, '(1X)' )
   WRITE (*,'( 1x,''sofa w/p03 vs w/o p03 ='',e20.10,'' uas'' )')
   . DROT ( R4, SOFTRANS ) *1D6 / AS2R

END

*----------------------------------------------------------------------*

SUBROUTINE REPMAT ( S, R )

* *
* ---PURPOSE:       Format and print 3x3 matrix
* *
* ---REFERENCES:    None
* *
* ---INPUT
* ARGUMENTS:        S = description of matrix
* (CHARACTER string)
* R = 3x3 matrix
  (DOUBLE PRECISION array)
*
*---OUTPUT
* ARGUMENTS: None
*
*---COMMON
* BLOCKS: None
*
*---ROUTINES
* CALLED: None
*
*---VER./DATE/
* PROGRAMMER: v1.0/02-09/AM (USNO/AA) initial version
*       v1.1/09-10/JLB (USNO/AA) additional documentation
*
-----------------------------------------------------------------------
IMPLICIT NONE
CHARACTER*(*) S
DOUBLE PRECISION R(3,3)
WRITE (*,'(/1X,A/SP,2(3F22.17/),3F22.17)') S,R(1,1),R(1,2),R(1,3),
:                   R(2,1),R(2,2),R(2,3),
:                   R(3,1),R(3,2),R(3,3)
C WRITE ( *, '(1X) ' )
C WRITE ( *,*) S
C WRITE ( *,'(3f22.17)') R
END

DOUBLE PRECISION FUNCTION DROT ( RMA, RMB )
*
*---PURPOSE: Express the difference between two rotation matrices
* RMA, RMB as an amount of rotation R about some
* arbitrary axis
*
*---REFERENCES: None
*
*---INPUT
* ARGUMENTS: RMA = First 3x3 matrix for comparison
*             (DOUBLE PRECISION array)
* RMA = Second 3x3 matrix for comparison
*             (DOUBLE PRECISION array)
*
*---OUTPUT
* ARGUMENTS: DROT = amount of rotation between RMA, RMB
*             (DOUBLE PRECISION)
*
*---COMMON
* BLOCKS: None
*
*---ROUTINES
* CALLED:    iau_TR   (SOFA)
*            iau_RXR  (SOFA)
*            iau_RM2V (SOFA)
*            iau_PM   (SOFA)
*
*---VER./DATE/
* PROGRAMMER: v1.0/02-09/AM  (USNO/AA) initial version
*              v1.1/09-10/JLB (USNO/AA) additional documentation
*
*-----------------------------------------------------------------------
IMPLICIT NONE
DOUBLE PRECISION RMA(3,3), RMB(3,3)

DOUBLE PRECISION RMBT(3,3), RM(3,3), RV(3), R

* Multiply the first matrix by the inverse of the second.
  CALL iau_TR ( RMB, RMBT )
  CALL iau_RXR ( RMBT, RMA, RM )

* Express the result as an r-vector.
  CALL iau_RM2V ( RM, RV )

* Return the magnitude (the amount of rotation in radians).
  CALL iau_PM ( RV, R )
  DROT = R

END

*----------------------------------------------------------------------*
SUBROUTINE TRANSPOSE (R, RT)
*
*
*---PURPOSE:     Transpose 3x3 matrix R ---> RT
*---REFERENCES:  None
*---INPUT
*   ARGUMENTS:    R = 3x3 matrix to be transposed
*                  (DOUBLE PRECISION array)
*---OUTPUT
*   ARGUMENTS:    RT = 3x3 matrix, transpose of R
*                  (DOUBLE PRECISION array)
*---COMMON
*   BLOCKS:       None
*---ROUTINES
*   CALLED:       None
*---VER./DATE/
* PROGRAMMER:  v1.0/02-09/AM  (USNO/AA) initial version
IMPLICIT NONE
DOUBLE PRECISION  R(3,3), RT(3,3)
INTEGER i, j

DO i = 1, 3
  DO j = 1, 3
    RT(i,j) = R(j,i)
  ENDDO
ENDDO

END

Input Data File: tercel_matrix
1  0.97310431770109063  0.23036382622411070 -0.00070316348296544
1 -0.23036380044102267  0.97310457063635536  0.00011854535854669
1  0.00071156016155651  0.00004662641201635  0.99999974575402495

Input Data File: sofa_matrix
1 +0.97310431757363003 +0.23036382624733387 -0.00070333224579995
1 -0.23036379880468646 +0.97310457073561185 +0.00012088727913663
1 +0.00071226387930021 +0.00004438635469672 +0.99999974535497649

Output Data
  0.973104317701091  0.230363826224111 -0.000703163482965
-0.230363800441023  0.973104570636355  0.000118545358547
  0.000711560161557  0.000046626412016  0.999999745754025

do 2007/04/05
time 12 0 0.0 UTC

X_p = +0.034928200 arcsec
Y_p = +0.483316300 arcsec
UT1-UTC = -0.072073685000 s

dDpsi (1980) = -0.055065500 arcsec
dDep (1980) = -0.006358000 arcsec
dX (2000) = +0.000172500 arcsec
dY (2000) = -0.000265000 arcsec
dX (2006) = +0.000175000 arcsec
dY (2006) = -0.000225900 arcsec

TT = 2400000.5 + 54195.500754444445192
UT1 = 2400000.5 + 54195.49999165811460

========================
4) IAU 2006/2000A, CIO-based

X =  +146.91514647359746
Y =    +9.155115079288397
s = -0.002200475 arcsec

NPB matrix, CIO-based (RC2I)
+0.99999974633944999 -0.00000000513882246 -0.00071226473007242
-0.00000002647522726  +0.999999991497483 -0.000004438524282713
+0.00071226472959891  +0.00004438525042571  +0.9999997453441982

DJMJD0, DATE, TUT:    2400000.5000000000        54195.0000000000
ERA = 0.2324515536471452 rad
= 13.318492965240 deg
= 00 53 16.438311658

celestial to terrestrial matrix (no polar motion)
+0.97310431757657001  +0.23036382623316559  -0.0007033281884719
-0.23036379873963870  +0.97310457073901657  +0.00012088854957536
+0.00071226472959891  +0.00004438525042571  +0.9999997453441982

celestial to terrestrial matrix
+0.97310431770097838  +0.23036382622458479  -0.00070316348220013
-0.23036380044149363  +0.97310457063624356  +0.0001185536661437
+0.00071156016266679  +0.00004662640399541  +0.99999974575402439

tercel_transposed matrix
+0.97310431770109063  +0.23036382622411070  -0.0007031634826544
-0.23036380044102267  +0.97310457063635536  +0.000118553654669
+0.00071156016155651  +0.00004662641201635  +0.99999974575402495

w/ pm result vs tercel = 0.1673896075E+01 uas
sofa w/p03 vs w/o p03 = 0.4843068713E+06 uas
Addendum III: C Version of SOFA Test Code

The C version of the SOFA test code consists of SOFA_Test.c and its accompanying header file, SOFA_Test.h. The SOFA Cookbook does not include examples in C. Therefore, SOFA_Test.c is basically a line-for-line transliteration of the Fortran SOFA-TEST.f that uses the SOFA C functions. At the end of the file, a C function terceltest, based on the Fortran program terceltest, is included.

SOFA_Test reads two input files: tercel-matrix.txt and sofa-matrix.txt. It produced the output reproduced at the end of this addendum.

terceltest reads a single input file: tercel-test-input.dat.

Program Header File: SOFA_Test.h

#include "stdio.h"
#include "sofa.h"
#include "novas.h"

void SOFA_Test ();
void repmat (char *s,double r[3][3]);
double drot (double rma[3][3],double rmb[3][3]);
void transpose (double r[3][3],double rt[3][3]);
void terceltest ();

Program File: SOFA_Test.c

#include "SOFA_Test.h"

void SOFA_Test ()
{
    /*  Arcseconds to radians*/
        double as2r = 4.848136811095359935899141e-6;
    /*  2Pi */
        double twopi = 6.283185307179586476925287;

    char pm;
    short year, month, day, hour, minute;
    int IHMSF[4];
    short i,j;

    double jd, seconds, xp, yp, dut1, ddp80, dde80, dx00, dy00, dx06, dy06,
            djmjd0, date, time, utc, dat, tai, TT, tut, ut1, result1,result2,
            rc2ti[3][3], rpm[3][3],rc2it[3][3], x, y, s, rc2i[3][3], era;
    double r3[3][3], r4[3][3];
    double termat[3][3], tertrans[3][3],sofmat[3][3], softrans[3][3];

    FILE *tercel_data,*sofa_data;

    tercel_data = fopen ("tercel_matrix.txt","r");
for (i=0; i<3; i++)
    for (j=0; j<3; j++)
        fscanf (tercel_data,"%lf", &termat[j][i]);
repmat ("termat", termat);

sofa_data = fopen ("sofa_matrix.txt","r");
for (i=0; i<3; i++)
    for (j=0; j<3; j++)
        fscanf (sofa_data,"%lf", &sofmat[j][i]);
repmat ("sofmat", sofmat);

/* utc. */
year = 2007;
month = 4;
day = 5;
hour = 12;
minute = 0;
seconds = 0.0;

printf ("date  %4i %2.2i  %2.2i\n", year, month, day);
printf ("time  %4i %2.2i  %5.1f UTC\n", hour, minute, seconds);

/* Polar motion (arcsec->radians). */
xp = 0.0349282 * as2r;
yp = 0.4833163 * as2r;
printf ("X_p =  %13.9f arcsec\n",xp/as2r);
printf ("Y_p =  %13.9f arcsec\n",yp/as2r);

/* ut1-utc (s). */
dut1 = -0.072073685;
printf ("ut1-utc = %16.12f\n", dut1);

ddp80 = -55.0655 * as2r / 1000.0;
dde80 = -6.3580 * as2r / 1000.0;
printf ("dDpsi (1980) = %13.9f arcsec\n",ddp80 / as2r);
printf ("dDepsi (1980) = %13.9f arcsec\n",dde80 / as2r);

/* CIP offsets wrt IAU 2000A (mas->radians). */
dx00 =  0.1725 * as2r / 1000.0;
dy00 = -0.2650 * as2r / 1000.0;
printf ("dx (2000) = %13.9f arcsec\n",dx00 / as2r);
printf ("dy (2000) = %13.9f arcsec\n",dy00 / as2r);

/* CIP offsets wrt IAU 2006/2000A (mas->radians). */
dx06 =  0.1750 * as2r / 1000.0;
dy06 = -0.2259 * as2r / 1000.0;
printf ("dx (2006) = %13.9f arcsec\n",dx06 / as2r);
printf ("dy (2006) = %13.9f arcsec\n",dy06 / as2r);

/* TT (MJD). */
iauCal2jd (year, month, day, &djmjd0, &date);
time = (60.0 * ((60.0 * (double) hour) + (double) minute) + seconds) / 86400.0;
utc = date + time;
iauDat (year, month, day, time, &dat);
tai = utc + dat / 86400.0;
TT = tai + 32.184 / 86400.0;
printf ("TT  = 2400000.5 + %22.15f\n", TT);
/*  ut1. */
tut = time + dut1 / 86400.0;
ut1 = date + tut;
printf ("UT1  = 2400000.5 + %22.15f\n", ut1);
/*  =========================
*  IAU 2006/2000A, CIO-based
*  ========================= */
printf ("=========================
IAU 2006/2000A, CIO-based
=========================
*  CIP and CIO, IAU 2006/2000A. */
/* Report inputs to CIP calculation */
printf ("djmjd0  = %22.15f tt= %22.15f\n", djmjd0,TT);
iauXys06a (djmjd0, TT, &x, &y, &s);
/*  Add CIP corrections. */
x += dx06;
y += dy06;
/*  Report CIP and s. */
printf ("x = %22.15f\n",x/as2r);
printf ("y = %22.15f\n",y/as2r);
printf ("s = %22.15f\n",s/as2r);
/*  GCRS to CIRS matrix. */
iauC2ixys (x, y, s, rc2i);
/*  Report. */
repmat ("NPB matrix, CIO-based (rc2i)", rc2i);
/*  Earth rotation angle. */
/*  Set tut to the value used in Tercel */
printf ("djmjd0 = %f  date = %f  tut = %f\n",djmjd0, date, tut);
era = iauEra00 (djmjd0, date+tut);
printf ("era = %19.16f rad\n",era);
printf ("era = %16.12f deg\n",era * 360.0/twopi);
iauD2tf (9, era/twopi, &pm, IHMSF);
/* Form celestial-terrestrial matrix (no polar motion yet). */
iauCr (rc2i, rc2ti);
iauRz (era, rc2ti);

/* Report. */
repmat ("celestial to terrestrial matrix (no polar motion)", rc2ti);
iauCr (rc2ti, r3);

/* Polar motion matrix (TIRS->ITRS, IERS 2003). */
iauPom00 (xp, yp, iauSp00(djmjd0, TT), rpom);

/* Form celestial-terrestrial matrix (including polar motion). */
iauRxr (rpom, rc2ti, rc2it);

/* Transpose termat */
transpose (termat, tertrans);
transpose (sofmat, softrans);

/* Report. */
repmat ("celestial to terrestrial matrix", rc2it);
repmat ("tercel_transposed matrix", tertrans);

/* Copy for later comparison. */
iauCr (rc2it, r4);

/* Compare result to TERCEL Transposed matrix */
result1 = drot (r4, tertrans);
printf ("%f\n", result1 / as2r);
printf ("w/ pm result vs tercel = %20.10e uas\n", drot (r4, tertrans) * 1.0e6 / as2r);

/* Compare SOFA w/P03 and SOFA w/o P03 */
result2 = drot (r4, softrans);
printf ("%f\n", result2 / as2r);
printf ("sofa w/p03 vs w/o p03 = %20.10e uas\n", drot (r4, softrans) * 1.0e6 / as2r);

}

/*@----------------------------------------------------------------------*/

void repmat (char *s, double r[3][3])
{
    printf ("%s\n", s);
    printf ("%22.17f %22.17f %22.17f\n", r[0][0], r[0][1], r[0][2]);
    printf ("%22.17f %22.17f %22.17f\n", r[1][0], r[1][1], r[1][2]);
    printf ("%22.17f %22.17f %22.17f\n", r[2][0], r[2][1], r[2][2]);
    return;
}
double drot (double rma[3][3], double rmb[3][3])
/* Express the difference between two rotation matrices RMA, RMB as an amount of rotation R about some arbitrary axis. */
{
    double rmbt[3][3], rm[3][3], rv[3], r;
    /* Multiply the first matrix by the inverse of the second. */
    iauTr (rmb, rmbt);
    iauRxr (rmbt, rma, rm);
    /* Express the result as an r-vector. */
    iauRm2v (rm, rv);
    /* Return the magnitude (the amount of rotation in radians). */
    r = iauPm (rv);
    return r;
}

void transpose (double r[3][3], double rt[3][3])
/*
Transpose 3x3 matrix R ---> RT
*/
{
    short i, j;
    for (i = 0; i<3; i++)
    {
        for (j = 0; j<3; j++)
            rt[i][j] = r[j][i];
    }
}

void terceltest ()
{
/* Transform vectors from ITRS to GCRS */
    double tjdh, tjd1, xp, yp, delt = 65.25607389, vec1[3], vec2[3], tjd,
dx, dy,mobl,tobl,ee;
    short num = 0;
    FILE *In_Data = NULL;
    dx = +0.1750;
    dy = -0.2259;
/* Open the input file of Julian dates, CIO coords, ITRS vector */

    In_Data = fopen("tercel-test-input.dat", "r");

    while (!feof(In_Data))
    {
        fscanf(In_Data,"%hi%lf%lf%lf%lf%lf%lf%lf", &num, &tjdh, &tjdl, &xp, &yp, &vec1[0], &vec1[1], &vec1[2]);

        /* Set transformation method, accuracy level, and ut1-utc. */
        tjd = tjdh + tjdl;

        /* celpol (tjd,2,dx,dy); */
        e_tilt (tjd, 0, &mobl, &tobl, &ee, &dx, &dy);

        /* Rotate vec1 from ITRS to GCRS = vec2 */
        ter2cel (tjdh, tjdl, delt, 1, 0, 0, xp, yp, vec1, vec2);

        printf("%i    %20.17f    %20.17f
%20.17f\n", num, vec2[0], vec2[1], vec2[2]);
    }

    fclose (In_Data);

Input Data File: tercel_matrix.txt
1 0.9731043170109063 0.23036382622411070 -0.00070316348296544
1 -0.23036380044102267 0.97310457063635536 0.00011854535854669
1 0.00071156016155651 0.00004662641201635 0.99999974575402495

Input Data File: sofa_matrix.txt
1 +0.97310431757363030 +0.23036382624733387 -0.0007033224579995
1 -0.23036379880468646 +0.97310457073561185 0.00012088727913663
1 +0.00071226387930021 +0.00004438635469672 0.99999974535497649

Input Data File: tercel-test-input.dat
1 2400000.5 54195.49999916581146 +0.0349282e0 +0.4833163e0 1.0 0.0 0.0
1 2400000.5 54195.49999916581146 +0.0349282e0 +0.4833163e0 0.0 1.0 0.0
1 2400000.5 54195.49999916581146 +0.0349282e0 +0.4833163e0 0.0 0.0 1.0

Output Data

termat
0.97310431770109063 -0.23036380044102267 0.00071156016155651
0.23036382622411070 0.97310457063635536 0.00004462641201635
-0.00070316348296544 0.00011854535854669 0.99999974575402495
sofmat
0.97310431757363030 -0.23036379880468646 0.00071226387930021
0.23036382624733387 0.97310457073561185 0.00004438635469672
date  2007 04  05
time    12 00    0.0 UTC

X_p =  0.034928200 arcsec
Y_p =  0.483316300 arcsec

ut1-utc =  -0.072073685000

dDpsi (1980) =  -0.055065500 arcsec
dDeps (1980) =  -0.006358000 arcsec
dx (2000) =   0.000172500 arcsec
dy (2000) =  -0.000265000 arcsec
dx (2006) =   0.000175000 arcsec
dy (2006) =  -0.000225900 arcsec

TT  = 2400000.5 +  54195.500754444445192
UT1  = 2400000.5 +  54195.499999165811460

=========================  
IAU 2006/2000A, CIO-based  
 =========================

x =    146.915146447359746
y =      9.155115079288397
s =      -0.002200474981084

NPB matrix, CIO-based (rc2i)
0.99999974633944499    -0.00000000513882246    -0.00071226473007242
-0.00000002647522726     0.99999999901497483    -0.00004438524282713
0.00071226472959891     0.00004438525042571     0.99999974633944499

djmjd0 = 2400000.500000  date = 54195.000000  tut = 0.499999

era =  0.2324515536471452 rad
era =  13.318492965240 deg

celestial to terrestrial matrix (no polar motion)
0.97310431757657001     0.23036382623316559    -0.00070333281884719
-0.23036379878963870     0.97310457073901657     0.00011854536661437
0.00071226472959891     0.00004438525042571     0.99999974633944499

celestial to terrestrial matrix
0.97310431770097838     0.23036382622458479    -0.0007031634820013
-0.23036380044149363     0.97310457063624356     0.00011854536661437
0.00071156016266793     0.00004662640399541     0.99999974575402495

tercel_transposed matrix
0.97310431770109063     0.23036382622411070    -0.00070316348296544
-0.23036380044102267     0.97310457063635536     0.00011854535854669
0.00071156016155651     0.00004662641201635     0.99999974575402495
w/ pm result vs tercel = $1.6738960753\times10^0$ uas

sofa w/p03 vs w/o p03 = $4.8430687135\times10^5$ uas